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Bose-Einstein Condensation for a Class of Homogeneous Interacting Systems

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Abstract

We prove rigorously the occurrence of zero-mode Bose-Einstein condensation for real continuous homogeneous systems of boson particles with superstable interactions and with a gap in the one-particle excitations spectrum. The idea of proof consists of comparing the system with a specially tuned soluble model.

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1 Introduction

Rigorous proofs of Bose-Einstein condensation (BEC) for realistic systems is an ongoing challenge with renewed interest after the successful experiments with trapped gases of alkali metals. In this note, we turn our attention to homogeneous continuous systems (i.e. without external trapping fields) of interacting bosons in the standard thermodynamic limit. This work started with the following underlying results in mind. First of all, the proof of Bose-Einstein condensation for systems with a gap in the one-particle excitation spectrum and with a Van der Waals family of two-body potentials by Buffet, de Smedt and Pulé [1]. The second is the proof of Bose-Einstein condensation in traps [2]. Although these two exact results are *a priori* unrelated, the first one being for homogeneous systems and the second one for inhomogeneous systems (because of the traps), both start from systems with a gap in the one-particle excitations spectrum. Heuristically, in both cases one might start from a system with a gap in the spectrum and with condensation in the groundstate. Contrary to systems without a gap, one can anticipate that the condensation for systems with a gap is stable under small perturbations. One might understand that condensation in the groundstate, which is energetically isolated by a gap, can survive the switching-on of a gentle interaction, and that fluctuations must be of microscopic size to overcome this gap and lift particles out of the isolated groundstate. This was our starting idea. First we define our systems under consideration.

Let us consider a gas of interacting bosons in hypercubic boxes $\Lambda = [-L/2, L/2]^\nu \subset \mathbb{R}^\nu$ of dimension $\nu \geq 1$ with periodic boundary conditions. The generalisation to other shapes for Λ is however readily done. Denote by $V = L^\nu$ the volume of the box Λ. The Hamiltonian for the volume Λ of this system in the boson Fock-space $\mathcal{F}_B$ reads

$$H^{\Delta}_{\Lambda,g} = T^{\Delta}_{\Lambda} - \mu N_\Lambda + gU_\Lambda, \qquad g > 0, \tag{1.1}$$

where T^{Δ}_{Λ} is the kinetic energy with the gap $\Delta > 0$ in its spectrum,

$$T^{\Delta}_{\Lambda} = \sum_{k\in\Lambda^*} \varepsilon_k a^\dagger_k a_k - \Delta a^\dagger_0 a_0, \qquad \varepsilon_k = \frac{\hbar^2 k^2}{2m}. \tag{1.2}$$

We take units $\hbar^2/2m = 1$, the sum k runs over the set Λ^*, dual to Λ, i.e.

$$\Lambda^* = \{k \in \mathbb{R}^\nu; k_\alpha = 2\pi n_\alpha/L; n_\alpha = 0, \pm 1, \ldots; \alpha = 1, 2, \ldots, \nu\}.$$

The operators $a^\dagger_k$ and a_k are the Bose creation and annihilation operators for mode $k \in \Lambda^*$, the number operators are denoted by $N_k = a^\dagger_k a_k$, and the total number operator in the volume Λ by $N_\Lambda = \sum_{k\in\Lambda^*} N_k$.

We assume *a priori* the presence of a gap Δ in the one-particle excitations spectrum, isolating the lowest energy level. This might seem rather artificial, however note that the presence of such a gap can be realised in several ways. A gap can be created using attractive boundary conditions [3–5], i.e. considering systems in containers Λ with *sticky* boundary conditions. But we have to notice that the Bose condensate is not homogeneous in this case. Another possibility is to assume that part of the particle interaction has caused this gap and is as such effectively incorporated [6]. The idea of considering a gap is not new. In his book [7] F. London attempted to introduce the gap on heuristic grounds to clarify some of the spectral properties of superfluid He^4.

The interaction between the particles is modelled by the two-body interaction operator

$$U_\Lambda = \frac{1}{2}\int_{\Lambda^2} \mathrm{d}x\mathrm{d}y\ a^\dagger(x)a^\dagger(y)v(x-y)a(y)a(x), \tag{1.3}$$

where $a^\dagger(x), a^\dagger(y)$ and $a(y), a(x)$ are the creation and annihilation operators for the Bose particles at $x, y \in \mathbb{R}^\nu$. Below we assume that the pair interaction potential $v(x)$ verifies the following conditions:

(a) $v : \mathbb{R}^\nu \to \mathbb{R}^1$ is a real, positive-type function from $L^1(\mathbb{R}^\nu)$. As it is known (the Bochner theorem [8]), a continuous function is of positive type if and only if it is the Fourier transform of a positive measure $\mathrm{d}\mu$ of finite total mass on $\mathbb{R}^\nu$:

$$v(x) = \int_{\mathbb{R}^\nu} \mathrm{d}\mu(x)\ \mathrm{e}^{\mathrm{i}qx} = v(-x).$$

(b) Since $v \in L^1(\mathbb{R}^\nu)$, the Fourier transform $\hat{v}(q)$ exists, and we suppose that

$$\hat{v}(0) = \int_{\mathbb{R}^\nu} \mathrm{d}x\ v(x) > 0 \quad \text{and,} \quad \hat{v}(0) \geq \hat{v}(q), \quad \forall q \in \mathbb{R}^\nu, \tag{1.4}$$

where $\hat{v}(q) \geq 0$ by (a).

Notice that by virtue of (a) the potential $v(x)$ at $x = 0$ is finite and that $v(0) \geq v(x)$, $\forall x \in \mathbb{R}^\nu$.

It is shown [9] that under conditions (a) and (b) the interaction is superstable, i.e. the n-body potential satisfies the inequality

$$\sum_{1\leq i<j\leq n} v(x_i - x_j) \geq \frac{A}{2V}n^2 - Bn$$

for some constants $A > 0$, $B \geq 0$, for all $n \in \mathbb{N}$, $x_i \in \Lambda$ and Λ large enough which implies that the thermodynamic potentials exist for all values of the chemical potential μ. Moreover, our proof requires a more stringent condition: it asks the constant $A = \hat{v}(0)(1-\epsilon)$, with $\epsilon > 0$ arbitrarily small and $B = v(0)/2$. These *optimal stability constants* for the L^1-integrable potentials of positive type were established by Lewis, Pulè and de Smedt [10]. Since these optimal constants are important for our proof, we incorporate in this note the corresponding argument (see section 2.1).

The superstability of the particle interaction is, together with the assumption of a spectral gap (1.2), the physical foundation of our proof. We prove that the zero-mode Bose-Einstein condensation in the groundstate of the one-particle spectrum isolated from the continuous spectrum by a gap Δ is robust enough under the application of a superstable interaction, provided that this spectral gap is large enough. Therefore, we show that in this case the zero-mode BEC is stable with respect to a class of superstable interactions. More technically our main result, announced in [11], can be stated as follows:

Theorem 1.1. *Consider a system of interacting Bose particles (1.1) in three or more dimensions, with a two-body interaction satisfying conditions (a) and (b). Fix an inverse temperature and a chemical potential* (β, μ) *such that* $\mu > g\hat{v}(0)\rho_c^P(\beta)$*, then there exists an non-empty interval of values of the gap* Δ*, such that for all* Δ *in this interval, the thermodynamic limit (* $\lim_\Lambda : V \to \infty$*) of the* $k = 0$ *mode particle number occupation density is positive:*

$$\rho_{0,g}^{\Delta}(\beta, \mu) \equiv \lim_{\Lambda} \frac{1}{V} \langle N_0 \rangle_{H_{\Lambda,g}^{\Delta}}(\beta, \mu) > 0.$$

Moreover, the Δ*-interval can be made arbitrarily large.* $\langle - \rangle_{H_{\Lambda,g}^{\Delta}}(\beta, \mu)$ *denotes the grand-canonical Gibbs state for the model* $H_{\Lambda,g}^{\Delta}$ *(1.1) at inverse temperature* $\beta = 1/k_B T$ *and chemical potential* μ*.* $\hat{v}(0)$ *is as in (1.4) and* $\rho_c^P(\beta)$ *is the critical density of the perfect Bose gas at inverse temperature* β*.*

A detailed proof is found in section 3, Figures 1–4 illustrating this theorem can be found in section 4. An analogous result as in Theorem 1.1 holds in lower dimensions $\nu \leq 2$ (cf. section 4). Our proof of BEC is based on comparing the condensate density of the full model (1.1) with the condensate density of specially tuned reference systems. This yields (various) lower bounds on the condensate of the full system in terms of thermodynamic potentials of the reference systems. The reference systems we use are based on the mean-field Bose gas. Its thermodynamic properties, necessary in our proof of BEC, are reviewed in section 2.2.

This paper is organised as follows: in section 2 the essential preliminaries, i.e. superstable potentials and the thermodynamics of the mean-field Bose gas, are reviewed, section 3 contains the actual proofs of the new results and a discussion on these results (section 4) concludes the paper.

2 Preliminaries

2.1 Superstable Potentials

An important aspect of particle interactions in both classical and quantum continuous systems are their stability properties [6, 9, 12]. If the particle interaction is such that the grand-canonical partition function does not converge, it is called catastrophic. It means that the thermodynamic pressure is not everywhere well defined, and good thermodynamic behaviour is excluded in those domains.

A useful criterion for stability is the following: a pair-potential v is called stable [9] if the corresponding n-particle interactions can be estimated from below as :

$$\sum_{1\leq i<j\leq n} v(x_i - x_j) \geq -Bn, \tag{2.1}$$

for a $B \geq 0$, and all $n \geq 2$, $x_i \in \Lambda$. This is a sufficient condition for good thermodynamic behaviour. An important subclass of the stable interactions are the so-called superstable pair-potentials [9], satisfying

$$\sum_{1\leq i<j\leq n} v(x_i - x_j) \geq \frac{A}{2V} n^2 - Bn \tag{2.2}$$

for some constants $A > 0$, $B \geq 0$, and all $n \geq 2$, $x_i \in \Lambda$, where Λ is large enough. This condition (2.2) yields the existence of the grand-canonical pressure for all values of the chemical potential $\mu \in \mathbb{R}$. From (2.2), it follows that the interaction term (1.3) satisfies the operator inequalty :

$$U_\Lambda \geq \frac{A}{2V} N_\Lambda^2 - BN_\Lambda. \tag{2.3}$$

An example of superstable potentials are those verifying the conditions (a) and (b). For this class of potentials Lewis, Pulè, and de Smedt [10] proved the existence of optimal constants $A = \hat{v}(0)(1-\epsilon)$ and $B = v(0)/2$ in (2.2). Here, $\epsilon > 0$ is an arbitrarily small positive constant. Since we use this result, we give a version of their proof adapted to our situation.

Lemma 2.1 (Lewis, Pulè, and de Smedt [10]). *Take $\epsilon > 0$, and take a real L^1-function of positive type $v : \mathbb{R}^\nu \to \mathbb{R}$ with $\hat{v}(0) > 0$ (1.4). There exist a subset $\Lambda_{min} \subset \mathbb{R}^\nu$ such that for each open box Λ, with $\Lambda_{min} \subset \Lambda$, the following inequality holds*

$$\sum_{1\leq i<j\leq n} v(x_i - x_j) \geq -\frac{v(0)}{2} n + \frac{\hat{v}(0)}{2V}(1-\epsilon)n^2, \tag{2.4}$$

for all $n \geq 2$, and each set of n distinct points $\{x_1, \ldots, x_n\} \subset \Lambda$.

Proof. By Bochner's theorem [8] a function of positive type v defines a positive-definite quadratic form $\langle\cdot,\cdot\rangle_v$ on the space of bounded measures on $\mathbb{R}^\nu$:

$$\langle \mu_1, \mu_2 \rangle_v = \iint_{\mathbb{R}^\nu \times \mathbb{R}^\nu} v(x-y)\ \mathrm{d}\mu_1(x)\mathrm{d}\mu_2(y).$$

Hence, by the Cauchy-Schwarz inequality this quadratic form satisfies

$$\langle \mu_1, \mu_1 \rangle_v \geq \frac{|\langle \mu_2, \mu_1 \rangle_v|^2}{\langle \mu_2, \mu_2 \rangle_v}.$$

Applying this inequality to the measures

$$\mathrm{d}\mu_1(x) = \sum_{i=1}^{n} \delta_{x_i}(x)\mathrm{d}x \text{ and } \mathrm{d}\mu_2(x) = \chi_{B(\Lambda,h)}(x)\mathrm{d}x,$$

where $\chi_{B(\Lambda,h)}(x)$ is the indicator of the set $B(\Lambda, h) = \{x \in \mathbb{R}^\nu : \mathrm{dist}(x, \Lambda) \leq h\}$, and $x_i \in \Lambda, i = 1, \ldots, n$, $n > 1$, we arrive at the following estimate:

$$\sum_{i,j=1}^{n} v(x_i - x_j) \geq \frac{\left|\sum_{i=1}^{n} A_\Lambda^h(x_i)\right|^2}{\int_{B(\Lambda,h)} \mathrm{d}y \ A_\Lambda^h(y)}, \tag{2.5}$$

where

$$A_\Lambda^h(y) = \int_{B(\Lambda,h)} \mathrm{d}x \ v(x-y).$$

A lower bound for the numerator in the r.h.s. of (2.5) is found using that for all $y \in \Lambda$ one has :

$$\begin{aligned} \left|A_\Lambda^h(y) - \hat{v}(0)\right| &\leq \left|A_\Lambda^h(y) - \int_{|x-y|\leq h} \mathrm{d}x \ v(x-y)\right| + \int_{|x|>h} \mathrm{d}x \ |v(x)| \\ &\leq \delta(h), \end{aligned} \tag{2.6}$$

where $\delta(h)$ is given by:

$$\delta(h) = 2\int_{|x|>h} \mathrm{d}x \ |v(x)|. \tag{2.7}$$

The last step in (2.6) is valid since for every $y \in \Lambda$, there is a ball with the radius h and the centre at y lying inside $B(\Lambda, h)$. Since v is an L^1-function, $\delta(h)$ (2.7) converges to zero in the limit $h \to \infty$. This yields

$$A_\Lambda^h(y) \geq \hat{v}(0) - \delta(h) > 0,$$

for all $y \in \Lambda$ and for h large enough. The last bound is valid since $\hat{v}(0) > 0$, and since $\delta(h) \geq 0$ (2.7) becomes arbitrarily small for h large enough. The numerator in the r.h.s. of (2.5) is therefore bounded from below as follows :

$$\left|\sum_{i=1}^{n} A_\Lambda^h(x_i)\right|^2 \geq n^2 \left(\hat{v}(0) - \delta(h)\right)^2. \tag{2.8}$$

An upper bound for the denominator in the r.h.s. of (2.5) is based on the following estimate

$$\begin{aligned} \left|\int_{B(\Lambda,h)} \mathrm{d}x \ A_\Lambda^h(y) - V\hat{v}(0)\right| &\leq \left|\int_{B(\Lambda,h)} \mathrm{d}y \ A_\Lambda^h(y) - \int_\Lambda \mathrm{d}y \ A_\Lambda^h(y)\right| + \left|\int_\Lambda \mathrm{d}y \ A_\Lambda^h(y) - V\hat{v}(0)\right| \\ &\leq \mathrm{vol}(B(\Lambda, h)\backslash\Lambda)\|v\|_{L^1} + V\delta(h), \end{aligned} \tag{2.9}$$

see (2.6). Hence, using the estimates (2.8) and (2.9) in (2.5), one gets

$$\sum_{i,j=1}^{n} v(x_i - x_j) \geq \frac{n^2}{V} \frac{(\hat{v}(0) - \delta(h))^2}{\hat{v}(0) + \|v\|_{L^1} \mathrm{vol}(B(\Lambda, h)\backslash\Lambda)/V + \delta(h)}.$$

Notice that for any smooth-shaped Λ the factor $\mathrm{vol}(B(\Lambda, h)\backslash\Lambda)/V$ is of the order $O(h/L)$, and vanishes in the limit $V \to \infty$ for fixed h. Since

$$\frac{(\hat{v}(0) - \delta(h))^2}{\hat{v}(0) + \|v\|_{L^1} \mathrm{vol}(B(\Lambda, h)\backslash\Lambda)/V + \delta(h)} \geq \hat{v}(0) - \|v\|_{L^1} \mathrm{vol}(B(\Lambda, h)\backslash\Lambda)/V - 3\delta(h),$$

we can choose h large enough, such that $\delta(h)/\hat{v}(0) < \epsilon/4$, and then take Λ_{min} that $\|v\|_{L^1} \mathrm{vol}(B(\Lambda, h)\backslash\Lambda)/V\hat{v}(0) < \epsilon/4$, to obtain the estimate :

$$\sum_{i,j=1}^{n} v(x_i - x_j) \geq \frac{n^2}{V} \hat{v}(0)(1 - \epsilon),$$

for all boxes Λ with $\Lambda_{min} \subset \Lambda$. □

As it follows from the proof, this result holds for more general shapes of Λ, the only condition is that $\mathrm{vol}(B(\Lambda, h)\backslash\Lambda)/V$ tends to zero in the limit $V \to \infty$ for some fixed $h > 0$. In fact, Λ tends to $\mathbb{R}^\nu$ in the sense of Van Hove [9] is enough.

In [10], it was also shown that the constants $A = \hat{v}(0)(1 - \epsilon)$ and $B = v(0)/2$ in Lemma 2.1 are optimal for this type of pair-potentials. This is based on the following argument: suppose there exists a series of positive constants $\{A_l\}_l$ converging for $l \to \infty$ to a better superstability constant $A > \hat{v}(0)$, i.e. such that for all l

$$\sum_{i,j} v(x_i - x_j) \geq \frac{n^2}{V_l} A_l, \tag{2.10}$$

for all finite sets of distinct points $\{x_1, \dots x_n\} \subset \Lambda_l$. Since we can choose $\epsilon > 0$ small enough that $A - \epsilon > \hat{v}(0)$, there exists $l_1(\epsilon)$ such that $A_l > A - \epsilon/2$ for all $l > l_1(\epsilon)$. On the other hand, for all Λ_l large enough, $l > l_2(\epsilon)$, we get

$$\int_{\Lambda_l} \mathrm{d}x \; v(x - y) \leq \hat{v}(0) + \epsilon/2,$$

uniformly in $y \in \mathbb{R}^\nu$. Then by integration of both sides of (2.10) over Λ_l^n, for $l > max(l_1(\epsilon), l_2(\epsilon))$ we get the estimates :

$$nv(0) + \frac{n(n-1)}{V_l}(\hat{v}(0) + \epsilon/2) \geq \frac{1}{V_l^n} \int_{{\Lambda_l}^n} \mathrm{d}x_1 \ldots \mathrm{d}x_n \; \sum_{i,j} v(x_i - x_j) \geq \frac{n^2}{V_l}(A - \epsilon/2).$$

Since above the $n > 1$ is arbitrary, these estimates imply that $A - \epsilon \leq \hat{v}(0)$, which is in contradiction to the hypothesis, hence yielding the optimality of the constants in Lemma 2.1.

2.2 Thermodynamics of the Mean-Field Bose Gas

In this section we briefly review the properties of the so-called mean-field Bose gas (sometimes also called the imperfect Bose gas), an exactly solvable model of Bosons [13–19] (for an extended review see [6]), which will play the rôle of a reference system. The mean-field Bose gas is defined by the local Hamiltonians

$$H^{\Delta}_{\Lambda,\lambda} = T^{\Delta}_{\Lambda} + \frac{\lambda}{2V} N^2_{\Lambda}. \tag{2.11}$$

The kinetic energy operator T^{Δ}_{Λ} (1.2) is identical to the one of the fully interacting system (1.1), but the interaction term (1.3) is replaced by a kind of mean-field interaction term. The physical relation of the model (2.11) to our system (1.1) lies in the fact that it is the Van der Waals limit of the fully interacting system (1.1) [1]. In that case, the constant λ has to be chosen equal to the long-range part of the interaction (1.3), i.e. $\lambda = g\hat{v}(0)$.

The explicit solution for the gapless case $\Delta = 0$ of this model can be found at several places [6, 13–19]. Here, we focus on the case with non-vanishing gap. The grand-canonical pressure function at inverse temperature β and chemical potential μ is defined as

$$p^{\Delta}_{\Lambda,\lambda}(\beta,\mu) = \frac{1}{\beta V} \ln \mathrm{Tr}_{\mathcal{F}_B}\, \mathrm{e}^{-\beta(H^{\Delta}_{\Lambda,\lambda} - \mu N_{\Lambda})}.$$

$\mathrm{Tr}_{\mathcal{F}_B}$ denotes the trace over the boson Fock space $\mathcal{F}_B$. Below, we develop explicit expressions for the pressure and the particle densities in the thermodynamic limit $V \to \infty$.

Lemma 2.2 (Thermodynamic Functions). *The grand-canonical pressure* $p^{\Delta}_{\lambda}(\beta,\mu) = \lim_{\Lambda} p^{\Delta}_{\Lambda,\lambda}(\beta,\mu)$ *of the mean-field Bose gas* (2.11) *exists for all* $\beta \geq 0, \mu \in \mathbb{R}$ *and is given by the Legendre transform:*

$$p^{\Delta}_{\lambda}(\beta,\mu) = \sup_{\rho \geq 0} \left(\mu\rho - f^{\Delta}_{\lambda}(\beta,\rho)\right), \tag{2.12}$$

where the canonical free energy $f^{\Delta}_{\lambda}(\beta,\rho)$ *at inverse temperature* β *and density* ρ *is given by*

$$f^{\Delta}_{\lambda}(\beta,\rho) = f^{P,\Delta}(\beta,\rho) + \frac{\lambda}{2}\rho^2, \tag{2.13}$$

$f^{P,\Delta}(\beta,\rho)$ *is the free energy of the perfect Bose gas with gap* Δ (1.2).

Proof. The thermodynamic pressure of the perfect Bose gas is given by:

$$p^{P,\Delta}(\beta,\mu) = \lim_{\Lambda} p^{P,\Delta}_{\Lambda}(\beta,\mu) = \lim_{\Lambda} \frac{1}{\beta V} \mathrm{Tr}_{\mathcal{F}_B}\, \mathrm{e}^{-\beta(T^{\Delta}_{\Lambda} - \mu N_{\Lambda})}$$

which implies that in order to be well defined, μ must be bounded from above: $\mu < -\Delta$, i.e.

$$p^{P,\Delta}_{\Lambda}(\beta,\mu) = \frac{1}{\beta V} \ln \sum_{n_0=0}^{\infty} \mathrm{e}^{\beta(\Delta+\mu)n_0} \sum_{\{n_k\}_{k\neq 0}} \mathrm{e}^{-\beta(\varepsilon_k - \mu)n_k}.$$

The canonical free energy $f^{P,\Delta}(\beta,\rho)$, is the Legendre transform of $p^{P,\Delta}(\beta,\mu)$, defined only for $\mu \leq -\Delta$,

$$f^{P,\Delta}(\beta,\rho) = \sup_{\mu\leq-\Delta} \left(\rho\mu - p^{P,\Delta}(\beta,\mu)\right). \tag{2.14}$$

By direct calculation one finds expression (2.13) for the free energy of the mean-field model (2.11) at temperature β and density ρ as

$$f_\lambda^\Delta(\beta,\rho) = \lim_\Lambda \frac{1}{\beta V} \ln \mathrm{Tr}_{\mathcal{H}_B^{(n)}} \mathrm{e}^{-\beta H_{\Lambda,\lambda}^\Delta},$$

where $\mathrm{Tr}_{\mathcal{H}_B^{(n)}}$ denotes the trace over the Hilbert space $\mathcal{H}_B^{(n)}$ of symmetrised functions for $n = \lfloor \rho V \rfloor$ Bosons. Since on this space the mean-field interaction term is constant, we immediately find :

$$\lim_\Lambda f_\Lambda[H_{\Lambda,\lambda}^\Delta](\beta,\rho) = \lim_\Lambda f_\Lambda[T_\Lambda^\Delta](\beta,\rho) + \frac{\lambda}{2}\rho^2.$$

The pressure of the mean-field gas, is again the Legendre transform of $f_\lambda^\Delta(\beta,\rho)$, yielding formula (2.12), well defined for all $\mu \in \mathbb{R}$. □

Theorem 2.3 (Pressure of the Mean-Field Bose Gas). *The grand-canonical pressure of the mean-field Bose Gas* (2.11) *is explicitly given by*

$$p_\lambda^\Delta(\beta,\mu) = \begin{cases} p_\lambda^{(\Delta=0)}(\beta,\mu), & \textit{for} \quad \mu \leq -\Delta + \lambda\rho^P(\beta,-\Delta); \\ (\mu+\Delta)^2/2\lambda + p^P(\beta,-\Delta), & \textit{for} \quad \mu > -\Delta + \lambda\rho^P(\beta,-\Delta), \end{cases} \tag{2.15}$$

where $p^P(\beta,\mu)$, and $\rho^P(\beta,\mu)$ are respectively the pressure, and the total density of the perfect Bose gas; $p_\lambda^{(\Delta=0)}(\beta,\mu)$ is the pressure of the mean-field Bose gas without gap.

Proof. The formula (2.15) is found by working out explicitly the Legendre transforms in (2.12)–(2.14), and using the properties of the perfect Bose gas. Let $\mu = \bar{\mu}_\Lambda(\beta,\rho)$ be solution of the equation:

$$\rho = \frac{1}{V}\langle N_\Lambda\rangle_{T_\Lambda^\Delta}(\beta,\bar{\mu}_\Lambda(\beta,\rho)),$$

for a given ρ and β, where the right-hand side is the expectation value of the total density in the grand-canonical Gibbs state for the perfect Bose gas model T_Λ^Δ. Denote the limiting solution by

$$\bar{\mu}(\beta,\rho) = \lim_\Lambda \bar{\mu}_\Lambda(\beta,\rho),$$

We have $\bar{\mu}(\beta,\rho) \leq -\Delta$, if $\rho \leq \rho_c^\Delta(\beta) \equiv \rho^P(\beta,-\Delta)$, and $\bar{\mu}(\beta,\rho) = -\Delta$, if $\rho \geq \rho^P(\beta,-\Delta)$, hence

$$f^{P,\Delta}(\beta,\rho) = \begin{cases} \rho\bar{\mu}(\beta,\rho) - p^P(\beta,\bar{\mu}), & \text{if} \quad \rho \leq \rho^P(\beta,-\Delta); \\ -\rho\Delta - p^P(\beta,-\Delta), & \text{if} \quad \rho > \rho^P(\beta,-\Delta). \end{cases} \tag{2.16}$$

This is the explicit expression for (2.14). The thermodynamic potentials such as the pressure and the particle density of the free Bose gas with gap Δ are the same as for the gapless perfect Bose gas, but only for the values of $\mu < -\Delta$. At $\mu = -\Delta$, there is degeneracy of the densities and BEC occurs. Since for $\Delta > 0$ the critical density $\rho_c^{P,\Delta}(\beta) \equiv \rho^P(\beta, -\Delta)$ is finite in all dimensions $\nu \geq 1$, the condensation takes place in all dimensions whereas in the gapless case $\rho_c^P(\beta) < \infty$ only for dimensions $\nu > 2$. Hence condensation occurs only in three or more dimensions at $\beta < \infty$.

Recall now the expression for the canonical free energy of the mean-field Bose gas, Eq. (2.13). Using the expression for the free energy of the perfect Bose gas $f^{P,\Delta}(\beta,\rho)$ (2.16) derived above, one finds for $\partial_\rho f_\lambda^\Delta(\beta,\rho)$,

$$\partial_\rho f_\lambda^\Delta(\beta,\rho) = \begin{cases} \partial_\rho f_\lambda^{P,\Delta}(\beta,\rho) + \lambda\rho, & \text{if} \quad \rho \leq \rho^P(\beta,-\Delta); \\ -\Delta + \lambda\rho, & \text{if} \quad \rho > \rho^P(\beta,-\Delta). \end{cases} \tag{2.17}$$

By virtue of (2.17) and (2.12), one gets the expression for the the mean-field Bose gas pressure:

$$\begin{aligned} p_\lambda^\Delta(\beta,\mu) &= \sup_{\rho\geq 0} \left(\mu\rho - f_\lambda^\Delta(\beta,\rho)\right) \\ &= \begin{cases} \mu\bar{\rho}(\beta,\mu) - f_\lambda^\Delta(\beta,\bar{\rho}(\beta,\mu)), & \text{for} \quad \mu \leq -\Delta + \lambda\rho^P(\beta,-\Delta); \\ \mu(\mu+\Delta)/\lambda - f_\lambda^\Delta(\beta,(\mu+\Delta)/\lambda), & \text{for} \quad \mu > -\Delta + \lambda\rho^P(\beta,-\Delta), \end{cases} \end{aligned}$$

where $\bar{\rho}(\beta,\mu)$ is the solution of $\partial_\rho f_\lambda^\Delta(\beta,\bar{\rho}(\beta,\mu)) + \lambda\bar{\rho}(\beta,\mu) = \mu$ as a function of $\mu \leq -\Delta + \lambda\rho^P(\beta,-\Delta)$ and β. Since by (2.16)

$$f^{P,\Delta}(\beta,(\mu+\Delta)/\lambda) = -\Delta\frac{\mu+\Delta}{\lambda} - p^P(\beta,-\Delta),$$

for $\mu + \Delta > \lambda\rho^P(\beta,-\Delta)$, where $p^P(\beta,-\Delta)$ is the pressure of the free Bose gas, we use the expression (2.13) for the free energy of the mean-field gas to find (2.15), that proves Theorem 2.3. □

Theorem 2.4. *Considering the mean-field Bose gas* (2.11)*, we derive the following expressions for the densities in the thermodynamic limit. The total grand-canonical density is given by*

$$\rho_\lambda^\Delta(\beta,\mu) = \begin{cases} \rho_\lambda^{(\Delta=0)}(\beta,\mu), & \textit{for} \quad \mu \leq -\Delta + \lambda\rho^P(\beta,-\Delta); \\ (\mu+\Delta)/\lambda, & \textit{for} \quad \mu > -\Delta + \lambda\rho^P(\beta,-\Delta). \end{cases} \tag{2.18}$$

The zero-mode condensate density is given by

$$\rho_{0,\lambda}^\Delta(\beta,\mu) = \begin{cases} 0, & \textit{for} \quad \mu \leq -\Delta + \lambda\rho^P(\beta,-\Delta); \\ (\mu+\Delta)/\lambda - \rho^P(\beta,-\Delta), & \textit{for} \quad \mu > -\Delta + \lambda\rho^P(\beta,-\Delta). \end{cases} \tag{2.19}$$

The limit of the expectation value $\langle N_\Lambda^2\rangle_{H_{\Lambda,\lambda}^\Delta}(\beta,\mu)/V^2$ is given by

$$\lim_\Lambda \frac{1}{V^2}\langle N_\Lambda^2\rangle_{H_{\Lambda,\lambda}^\Delta}(\beta,\mu) = \begin{cases} \rho_\lambda^{(\Delta=0)}(\beta,\mu)^2, & \textit{for} \quad \mu \leq -\Delta + \lambda\rho^P(\beta,-\Delta); \\ (\mu+\Delta)^2/\lambda^2, & \textit{for} \quad \mu > -\Delta + \lambda\rho^P(\beta,-\Delta). \end{cases} \tag{2.20}$$

Here $\rho_\lambda^{(\Delta=0)}(\beta,\mu)$ is the density for the gapless mean-field gas, i.e. for $\Delta = 0$ in Eq. (2.11).

Proof. These quantities are derived using that the pressure (2.15) is a convex function of respectively Δ, μ and λ. By the Griffith lemma [6, Appendix C], the order of the thermodynamic limit and the corresponding derivative can be interchanged, which gives :

$$\rho^{\Delta}_{0,\lambda}(\beta,\mu) = \lim_{\Lambda} \frac{1}{V}\langle N_0\rangle_{H^{\Delta}_{\Lambda,\lambda}}(\beta,\mu) = \lim_{\Lambda} \partial_{\Delta} p^{\Delta}_{\Lambda,\lambda}(\beta,\mu) = \partial_{\Delta} p^{\Delta}_{\lambda}(\beta,\mu);$$

$$\rho^{\Delta}_{\lambda}(\beta,\mu) = \lim_{\Lambda} \frac{1}{V}\langle N_\Lambda\rangle_{H^{\Delta}_{\Lambda,\lambda}}(\beta,\mu) = \lim_{\Lambda} \partial_{\mu} p^{\Delta}_{\Lambda,\lambda}(\beta,\mu) = \partial_{\mu} p^{\Delta}_{\lambda}(\beta,\mu);$$

$$\lim_{\Lambda} \frac{1}{V^2}\langle N^2_\Lambda\rangle_{H^{\Delta}_{\Lambda,\lambda}}(\beta,\mu) = \lim_{\Lambda} -2\partial_{\lambda} p^{\Delta}_{\Lambda,\lambda}(\beta,\mu) = -2\partial_{\lambda} p^{\Delta}_{\lambda}(\beta,\mu).$$

By virtue of (2.15) of Theorem 2.3 these imply the explicit expressions (2.18)–(2.20) of Theorem 2.4. □

Taking the limit $\Delta \downarrow 0$, we recover the usual expressions for the mean-field Bose gas (2.11) with vanishing gap, in particular the expression for the zero-mode condensate density in dimensions $\nu \geq 3$,

$$\rho^{(\Delta=0)}_{0,\lambda}(\beta,\mu) = \begin{cases} 0, & \text{for } \mu \leq \lambda\rho^P_c(\beta); \\ \mu/\lambda - \rho^P_c(\beta), & \text{for } \mu > \lambda\rho^P_c(\beta). \end{cases} \tag{2.21}$$

3 Proofs of the Main Results

The main idea of the proof of the condensation for the systems (1.1) is to estimate their Bose condensate from below by the condensate of a particularly chosen reference system for which one can compute the amount of the condensate explicitly. Thus, a judicious choice of this reference system is a subtle point of our proof.

Since we consider superstable systems, i.e. systems where the grand-canonical pressure is defined for all values of the chemical potential, it seems to be natural to choose a reference system which is also superstable. This immediately rules out the perfect Bose gas (1.2) as a reference system, since its pressure is only well defined for $\mu \leq -\Delta$. Choosing the reference systems within the class of mean-field Bose gases (cf. section 2.2), which are indeed well-known superstable systems, seems therefore a good choice. The reference systems that we consider are mean-field Bose systems which are *close enough* to the Van der Waals limit of the fully interacting system (1.1). It turns out that the Van der Waals limit itself is *not* the best choice. Apart from the use of a reference system, the proof is based on various convexity properties of the thermodynamic functions. In particular it is based on the following lemma.

Lemma 3.1. *The zero-mode condensate density $\rho^{\Delta}_{0,g}(\beta,\mu)$ in the thermodynamic limit of grand-canonical Gibbs states of interacting system (1.1) with a superstable two-body* *potentials v satisfying the conditions (a) and (b), has the following lower bound :*

$$\begin{aligned}\rho_{0,g}^{\Delta}(\beta,\mu) \geq \frac{\mu}{g\hat{v}(0)(1-\epsilon)} + \frac{g\hat{v}(0)}{2\Delta}\rho^{P}(\beta,-\Delta)^{2} - \frac{gv(0)}{2\Delta}\rho_{g}^{(\Delta=0)}(\beta,\mu) \\ - \frac{(\mu+\Delta)^{2}}{2\Delta g\hat{v}(0)}\frac{\epsilon}{(1-\epsilon)^{2}} - \frac{\mu+\Delta}{\Delta(1-\epsilon)}\rho^{P}(\beta,-\Delta) - \rho_{c}^{P}(\beta).\end{aligned} \tag{3.1}$$

Here $\rho_{g}^{(\Delta=0)}(\beta,\mu)$ denotes the total density of the interacting gas without gap (1.1), and $\epsilon > 0$ is an arbitrarily small constant. $\rho^{P}(\beta,-\Delta)$ refers to the total density of the perfect Bose gas at the inverse temperature β and the chemical potential $\mu = -\Delta$, $\rho_{c}^{P}(\beta)$ is the critical density of the perfect Bose gas. The bound is valid for values $\mu > g\hat{v}(0)(1-\epsilon)\rho_{c}^{P}(\beta)$, and dimensions $\nu \geq 3$.

Proof. The pressure $p_{\Lambda}[H_{\Lambda}^{\Delta}]$ of systems with a gap in the kinetic energy spectrum (1.2) and any stable interaction is an increasing convex function of the parameter $\Delta \geq 0$. Since by Theorem 2.4 the condensate density $\rho_{0,g}^{\Delta}(\beta,\mu)$ is the derivative of the corresponding pressure with respect to Δ, the convexity property yields a lower bound for the zero-mode density $\langle N_0/V\rangle_{H_{\Lambda,g}^{\Delta}}$:

$$\frac{1}{V}\langle N_0\rangle_{H_{\Lambda,g}^{\Delta}} \geq \frac{p_{\Lambda}[H_{\Lambda,g}^{\Delta}] - p_{\Lambda}[H_{\Lambda,g}^{(\Delta=0)}]}{\Delta}. \tag{3.2}$$

Now we use a reference system to get the lower bound on the condensate. This reference system is a mean-field Bose gas (2.11), defined by the local Hamiltonian (cf. section 2.2)

$$H_{\Lambda,g,A}^{\Delta} = T_{\Lambda}^{\Delta} - \mu N_{\Lambda} + g\frac{A}{2V}N_{\Lambda}^{2}. \tag{3.3}$$

Therefore, we fix the mean-field Bose gas (2.11) interaction parameter by taking $\lambda = gA$, where $g > 0$ is the coupling constant (cf. (1.1)) and $A = \hat{v}(0)(1-\epsilon)$ is the optimal superstability constant (2.2) associated with the two-body interaction (1.3) of the full model (1.1). By virtue of the same convexity property as in (3.2), the difference of the pressures between the reference Bose system (3.3) with gap and without gap, is bounded from below by the condensate density for the reference mean-field gas without gap, i.e.

$$\frac{p_{\Lambda}[H_{\Lambda,g,A}^{\Delta}] - p_{\Lambda}[H_{\Lambda,g,A}^{(\Delta=0)}]}{\Delta} \geq \frac{1}{V}\langle N_0\rangle_{H_{\Lambda,g,A}^{(\Delta=0)}}. \tag{3.4}$$

Adding the inequality (3.4) to the lower bound on the condensate density of the full system (3.2), we introduce the reference system (3.3) in our estimate:

$$\frac{1}{V}\langle N_0\rangle_{H_{\Lambda,g}^{\Delta}} \geq \frac{1}{V}\langle N_0\rangle_{H_{\Lambda,g,A}^{(\Delta=0)}} - \frac{1}{\Delta}\left(p_{\Lambda}[H_{\Lambda,g,A}^{\Delta}] - p_{\Lambda}[H_{\Lambda,g,A}^{(\Delta=0)}] - p_{\Lambda}[H_{\Lambda,g}^{\Delta}] + p_{\Lambda}[H_{\Lambda,g}^{(\Delta=0)}]\right). \tag{3.5}$$

Hence, the condensate density of the interacting model (1.1) with gap ($\Delta > 0$) is bounded from below by the condensate density of the mean-field model (3.3) without gap ($\Delta = 0$), and a correction term proportional to $1/\Delta$ containing the pressure differences between the full system and the reference system.

These pressure differences will be estimated using the Bogoliubov convexity inequality [6, Appendix D]. Applied to the grand-canonical pressures of the mean-field reference Bose gas (3.3) and the full model (1.1), it gives

$$\frac{g}{V}\langle W_\Lambda^A\rangle_{H_{\Lambda,g}^\Delta} \leq p_\Lambda[H_{\Lambda,g,A}^\Delta] - p_\Lambda[H_{\Lambda,g}^\Delta] \leq \frac{g}{V}\langle W_\Lambda^A\rangle_{H_{\Lambda,g,A}^\Delta}, \tag{3.6}$$

for any $\Delta \geq 0$. Here the operator W_Λ^A is the difference between the interactions of the fully interacting and the mean-field Bose gases : $W_\Lambda^A = U_\Lambda - AN_\Lambda^2/2V$. Then by virtue of (3.5) and (3.6) we get :

$$\frac{1}{V}\langle N_0\rangle_{H_{\Lambda,g}^\Delta} \geq \frac{1}{V}\langle N_0\rangle_{H_{\Lambda,g,A}^{(\Delta=0)}} - \frac{g}{\Delta}\left(\frac{1}{V}\langle W_\Lambda^A\rangle_{H_{\Lambda,g,A}^\Delta} - \frac{1}{V}\langle W_\Lambda^A\rangle_{H_{\Lambda,g}^{(\Delta=0)}}\right). \tag{3.7}$$

Now our task is to estimate the two expectation values of W_Λ^A in (3.7). An upper bound on $\langle W_\Lambda^A/V\rangle_{H_{\Lambda,g,A}^\Delta}$ in (3.7) can be found using the properties of the pair-potential v and the Gibbs states of the reference system (3.3). Expressed in terms of the creation and annihilation operators on Λ^*, we get for $\langle W_\Lambda^A/V\rangle_{H_{\Lambda,g,A}^\Delta}$:

$$\begin{aligned}\frac{1}{V}\langle W_\Lambda^A\rangle_{H_{\Lambda,g,A}^\Delta} = {} & \frac{1}{2V^2}\sum_{q\in\Lambda^*}\sum_{k\in\Lambda^*}\sum_{k'\in\Lambda^*}\hat{v}(q)\langle a_{k'+q}^\dagger a_{k-q}^\dagger a_k a_{k'}\rangle_{H_{\Lambda,g,A}^\Delta} \\ & - \frac{1}{2V^2}A\sum_{k\in\Lambda^*}\sum_{k'\in\Lambda^*}\langle a_{k'}^\dagger a_{k'} a_k^\dagger a_k\rangle_{H_{\Lambda,g,A}^\Delta}.\end{aligned}$$

Exploiting the mode by mode gauge invariance of the Gibbs states of the mean-field Bose gas (3.3), and rewriting the above expression in terms of the occupation-number operators $N_k = a_k^\dagger a_k$ we arrive at

$$\begin{aligned}\frac{1}{V}\langle W_\Lambda^A\rangle_{H_{\Lambda,g,A}^\Delta} = {} & \frac{1}{2V^2}\sum_k\sum_{k'}\left(\hat{v}(0)+\hat{v}(k-k')-A\right)\langle N_k N_{k'}\rangle_{H_{\Lambda,g,A}^\Delta} \\ & - \frac{1}{2V^2}\hat{v}(0)\sum_k\left(\langle N_k^2\rangle_{H_{\Lambda,g,A}^\Delta} + \langle N_k\rangle_{H_{\Lambda,g,A}^\Delta}\right).\end{aligned} \tag{3.8}$$

Since by condition (b) $\hat{v}(0) \geq \hat{v}(k) \geq 0$, the coefficients in the first sum of the r.h.s. of (3.8) are bounded as

$$\frac{1}{2}\left(\hat{v}(0)+\hat{v}(k-k')-A\right) \leq \hat{v}(0) - A/2.$$

From the second sum in the r.h.s. of (3.8), we retain only the quadratic zero-mode term contained in

$$-\frac{\hat{v}(0)}{2V^2}\langle N_0^2\rangle_{H_{\Lambda,g,A}^\Delta} = -\frac{\hat{v}(0)}{2V^2}\left(\langle N_0\rangle_{H_{\Lambda,g,A}^\Delta}^2 + \langle N_0\rangle_{H_{\Lambda,g,A}^\Delta}\right).$$

This yields the following upper bound for $\langle W_\Lambda^A/V\rangle_{H^\Delta_{\Lambda,g,A}}$:

$$\frac{1}{V}\langle W_\Lambda^A\rangle_{H^\Delta_{\Lambda,g,A}} \leq \frac{2\hat{v}(0)-A}{2V^2}\langle N_\Lambda^2\rangle_{H^\Delta_{\Lambda,g,A}} - \frac{\hat{v}(0)}{2V^2}\langle N_0\rangle^2_{H^\Delta_{\Lambda,g,A}}. \tag{3.9}$$

The expectation values appearing in the r.h.s. of (3.9) can be calculated exactly, applying Theorem 2.4. They give in the thermodynamic limit the upper bound : $(\hat{v}(0) - A/2)\rho^\Delta_{g,A}(\beta,\mu)^2 - \hat{v}(0)\rho^\Delta_{0,g,A}(\beta,\mu)^2/2$.

The other unknown term in (3.7) is $\langle W_\Lambda^A\rangle_{H^{(\Delta=0)}_{\Lambda,g}}$. It can be estimated using the superstability (2.3) of the interaction U_Λ (1.3) by the tuning the interaction parameter of the mean-field reference Bose gas (3.3) to be equal to the constant A in the superstability criterion (2.3), which gives the estimate from below :

$$\frac{1}{V}\langle W_\Lambda^A\rangle_{H^{(\Delta=0)}_{\Lambda,g}} \geq -\frac{B}{V}\langle N_\Lambda\rangle_{H^{(\Delta=0)}_{\Lambda,g}}. \tag{3.10}$$

This, in particular, justifies our choice of the parameter $\lambda = gA$ specifying the reference system (3.3). Using now (3.9) and (3.10) in (3.7) one finds in the thermodynamic limit the following lower bound for the condensate density $\rho^\Delta_{0,g}(\beta,\mu)$

$$\begin{aligned}\rho^\Delta_{0,g}(\beta,\mu) \geq \rho^{(\Delta=0)}_{0,g,A}(\beta,\mu) + g\frac{\hat{v}(0)}{2\Delta}\rho^\Delta_{0,g,A}(\beta,\mu)^2 \\ - \frac{g}{\Delta}\left(B\rho_g^{(\Delta=0)}(\beta,\mu) + (\hat{v}(0)-A/2)\rho^\Delta_{g,A}(\beta,\mu)^2\right).\end{aligned} \tag{3.11}$$

The lower bound (3.1) now follows from the explicit expressions (Theorem 2.4) for the total density and the condensate density of the mean-field Bose gas with gap, and from the well-known expression for the condensate density in the gapless mean-field model (2.21) for $\mu > gA\rho_c^P(\beta)$, i.e. in the regime where $\rho^{(\Delta=0)}_{0,g,A} > 0$.

In the last step to (3.1) we use the optimal superstability constants for continuous L^1-potentials of positive type (cf. Lemma 2.1), i.e. we put $A = (1-\epsilon)\hat{v}(0)$, and $B = v(0)/2$. This gives the expression for the lower bound in the the form (3.1). □

Notice that the lower bound (3.1) contains the term $\rho_g^{(\Delta=0)}(\beta,\mu)$, i.e. the total density of the fully interacting gas without gap. It is not explicitly known as a function of β and μ. However it is always finite, and it can be viewed as a reference parameter. Using a slightly modified reference system, an alternative lower bound can be derived which consists only of explicitly know functions related to the perfect Bose gas.

Lemma 3.2. *The zero-mode condensate density $\rho^\Delta_{0,g}(\beta,\mu)$ in the thermodynamic limit of the grand-canonical Gibbs states of interacting systems (1.1) with superstable two-body potential v satisfying conditions (a) and (b), has the following alternative lower bound:*

$$\begin{aligned}\rho^\Delta_{0,g}(\beta,\mu) \geq \frac{2\mu+gv(0)}{2g\hat{v}(0)(1-\epsilon)} + \frac{g\hat{v}(0)}{2\Delta}\rho^P(\beta,-\Delta)^2 - \frac{(2\mu+2\Delta+gv(0))^2}{8g\Delta\hat{v}(0)}\frac{\epsilon}{(1-\epsilon)^2} \\ - \frac{2\mu+2\Delta+gv(0)}{2\Delta\hat{v}(0)(1-\epsilon)}\left(\frac{v(0)}{2} + \hat{v}(0)\rho^P(\beta,-\Delta)\right) - \rho_c^P(\beta).\end{aligned} \tag{3.12}$$

Here $\epsilon > 0$ is an arbitrarily small constant, $\rho^P(\beta, -\Delta)$ refers to the total density of the perfect Bose gas at inverse temperature β and chemical potential $\mu = -\Delta$, and $\rho_c^P(\beta)$ is the critical density of the perfect Bose gas. The bound is valid for all values $\mu > g\hat{v}(0)(1-\epsilon)\rho_c^P(\beta)$, and dimensions $\nu \geq 3$.

Proof. The proof is completely analogous to the proof of Lemma 3.1. But now we use the alternative reference system :

$$H^{\Delta}_{\Lambda,g,A,B} = T^{\Delta}_{\Lambda} - \mu N_{\Lambda} + g\left(\frac{A}{2V}N^2_{\Lambda} - BN_{\Lambda}\right), \tag{3.13}$$

which compared to the first reference system (3.3), contains an extra interaction term. Since the term $-gBN_{\Lambda}$ is linear in the total number operator, it corresponds to a shift in the chemical potential. Again, the constants A and B coincide with the optimal superstability values (2.2) for the pair-potential v of the full system (1.1), where $g > 0$.

First, we derive a bound similar to the one of (3.7) in the proof of Lemma 3.1. Now one gets :

$$\frac{1}{V}\langle N_0\rangle_{H^{\Delta}_{\Lambda,g}} \geq \frac{1}{V}\langle N_0\rangle_{H^{(\Delta=0)}_{\Lambda,g,A,B}} - \frac{g}{\Delta}\left(\frac{1}{V}\langle W^{A,B}_{\Lambda}\rangle_{H^{\Delta}_{\Lambda,g,A,B}} - \frac{1}{V}\langle W^{A,B}_{\Lambda}\rangle_{H^{(\Delta=0)}_{\Lambda,g}}\right), \tag{3.14}$$

where $W^{A,B}_{\Lambda} = U_{\Lambda} - AN^2_{\Lambda}/2V + BN_{\Lambda}$. The expectation values in the r.h.s. of (3.14) can be estimated analogously to (3.9) and (3.10). This yields for the upper bound :

$$\frac{1}{V}\langle W^{A,B}_{\Lambda}\rangle_{H^{\Delta}_{\Lambda,g,A,B}} \leq \frac{2\hat{v}(0) - A}{2V^2}\langle N^2_{\Lambda}\rangle_{H^{\Delta}_{\Lambda,g,A,B}} - \frac{\hat{v}(0)}{2V^2}\langle N_0\rangle^2_{H^{\Delta}_{\Lambda,g,A,B}} + \frac{B}{V}\langle N_{\Lambda}\rangle_{H^{\Delta}_{\Lambda,g,A,B}}. \tag{3.15}$$

For the lower bound of $\langle W^{A,B}_{\Lambda}\rangle_{H^{(\Delta=0)}_{\Lambda,g}}$ we use again the superstability of the interaction U_{Λ} (1.3), and the fact that according to the superstability criterion (2.3) we can take for A and B their optimal values. This gives:

$$\frac{1}{V}\langle W^{A,B}_{\Lambda}\rangle_{H^{(\Delta=0)}_{\Lambda,g}} \geq 0. \tag{3.16}$$

The explicit formula (3.12) now follows if one introduces (3.15) and (3.16) into (3.14), using the explicit expressions for the densities of the mean-field Bose gas (section 2.2), and for the optimal values of A and B (section 2.1), and finally taking the thermodynamic limit. □

It should be remarked that one can hardly compare the bound given in Lemma 3.1 with the one in Lemma 3.2, and hence to express an opinion which of them yields the best result. However, as the latter bound is known explicitly, it can be used to make numerical estimates of the condensate density and of the minimal gap as functions of the various parameters involved. For this we refer to ref. [11] and section 4.

Instead, we proceed now with the proof of our Theorem 1.1, based on the lower bound derived in Lemma 3.1. We prove that the condensate density of the full model (1.1) is strictly positive in the domain $\mu > g\hat{v}(0)\rho_c^P(\beta)$ if the gap is large enough.

Proof of Theorem 1.1. It follows from Lemma 3.1 that for every $\epsilon < 1$ and any $\epsilon' < 1$ such that $\epsilon' > \epsilon > 0$ one gets by (3.1) the estimate:

$$\begin{aligned}\rho_{0,g}^{\Delta}(\beta,\mu) \geq \; & \frac{\mu}{g\hat{v}(0)} - \rho_c^P(\beta) \\ & + \frac{g\hat{v}(0)}{2\Delta}\rho^P(\beta,-\Delta)^2 - \frac{gv(0)}{2\Delta}\rho_g^{(\Delta=0)}(\beta,\mu) - \frac{\mu+\Delta}{\Delta(1-\epsilon')}\rho^P(\beta,-\Delta) \\ & - \frac{(\mu+\Delta)^2}{2\Delta g\hat{v}(0)}\frac{\epsilon}{(1-\epsilon)^2}.\end{aligned} \tag{3.17}$$

Fix the inverse temperature β and take the chemical potential μ compatible with the condition of the theorem. This ensures the positivity of the first term in the r.h.s. of (3.11)). Now take μ such that

$$\mu/g\hat{v}(0) - \rho_c^P(\beta) > 3\eta,$$

for some arbitrarily chosen $\eta > 0$. This yields a lower bound for the first line in the r.h.s. of (3.17).

The expression on second line can be absolutely bounded by η using Δ large enough and the fact that $\lim_{\Delta\to\infty}\rho^P(\beta,-\Delta) = 0$. This gives :

$$\left|\frac{g\hat{v}(0)}{2\Delta}\rho^P(\beta,-\Delta)^2 - \frac{gv(0)}{2\Delta}\rho_g^{(\Delta=0)}(\beta,\mu) - \frac{\mu+\Delta}{\Delta(1-\epsilon')}\rho^P(\beta,-\Delta)\right| < \eta,$$

for all Δ larger than some minimal gap: $\Delta \geq \tilde{\Delta}(\mu,\epsilon',\eta)$, which exists for the fixed $\mu > g\hat{v}(0)\rho_c^P(\beta)$ and $\epsilon' < 1$.

Finally for the third line in (3.17), we can choose ϵ small enough to ensure the estimate

$$\frac{(\mu+\Delta)^2}{2\Delta g\hat{v}(0)}\frac{\epsilon}{(1-\epsilon)^2} < \eta,$$

in the interval $[\Delta_1,\Delta_2]$ with $\tilde{\Delta}(\mu,\epsilon',\eta) < \Delta_2$, see Figure 1. Collecting all these estimates, we obtain that for a fixed temperature and $\eta > 0$ one can find μ large enough and ϵ small enough that they ensure the existence of a non-empty interval of Δ such that :

$$\rho_{0,g}^{\Delta}(\beta,\mu) > \eta > 0,$$

in this interval, proving the condensation. The set of allowed values of Δ is confined by the estimate $\tilde{\Delta}(\mu,\epsilon',\eta) < \Delta_2$. Since $\Delta_2 = \Delta_2(\epsilon)$ is increasing when ϵ gets smaller, the non-empty Δ-interval exists for all other parameters fixed. □

Similar one can prove the existence of the zero-mode condensation on the basis of the bound found in Lemma 3.2.

4 Discussion

So far, we are concentrated on the case of three or more dimensions $\nu \geq 3$, however, the result of Theorem 1.1 can be extended to dimensions $\nu = 1$ or $\nu = 2$. A lower bound for the condensate density $\rho_{0,g}^{\Delta}(\beta, \mu)$ for $\nu \leq 2$ as in Lemma 3.1 or Lemma 3.2 is derived in a similar way. It requires slightly modified convexity arguments (3.2)–(3.4). Since the free Bose gas in dimensions $\nu \leq 2$ shows only condensation in the case of non-vanishing gap (1.2), one has to consider in (3.2)–(3.4) the pressure differences in the form $p_\Lambda[H_\Lambda^{\Delta}] - p_\Lambda[H_\Lambda^{\Delta_0}]$, for some $\Delta_0 > 0$, with $0 < \Delta_0 < \Delta$, instead of $\Delta_0 = 0$. This yields the substitution in (3.11) of $\rho_{0,g,A}^{(\Delta=0)}(\beta, \mu)$ and $\rho_g^{(\Delta=0)}(\beta, \mu)$ by $\rho_{0,g,A}^{\Delta_0}(\beta, \mu)$ and $\rho_g^{\Delta_0}(\beta, \mu)$. The bounds derived in this way are valid for all dimensions $\nu \geq 1$, and lead to similar conclusions as in Theorem 1.1. Hence, in one and two dimensional interacting Bose gases with large enough gap (1.2), the zero-mode Bose-Einstein condensation is also proved. Notice that this is in contrast to the Bogoliubov-Hohenberg theorem [12, 20] which yields the absence of BEC for translation invariant continuous Bose systems without gap for dimensions $\nu \leq 2$.

We use the lower bound (3.12), which can be computed explicitly as a function of the different parameters, to visualise our estimates. For example, the Figure 1 shows the ϵ-dependence of the condensation domain. Two solid lines shown on the (μ, Δ)-graph

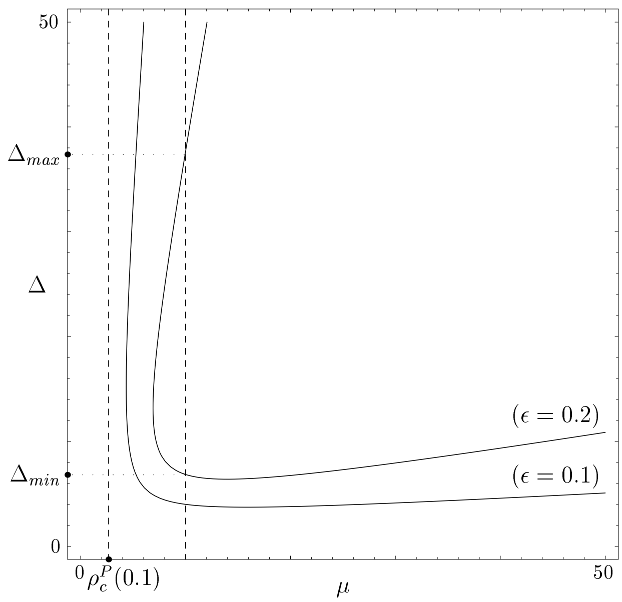

Figure 1: (μ, Δ)-graph with ϵ-dependence

indicate the borders of domains where the lower bound (3.12) is positive, i.e. in the domain surrounded by these curves we have BEC. The lowest line is computed for $\epsilon = 0.1$, whereas the upper line is computed for $\epsilon = 0.2$, both lines are for $\beta = 0.1$, and $g = \hat{v}(0) = v(0) = 1$. The left dashed vertical line indicates the critical value for the chemical potential $\mu = g\hat{v}(0)\rho_c^P(\beta)$ above which (according to Theorem 1.1) our bounds are valid.

On this figure we also illustrate the line of reasoning of Theorem 1.1: fixing a temperature, and choosing a chemical potential μ larger than the critical value $\mu = \rho_c^P(0.1)$, we find, for a small enough ϵ, an interval $[\Delta_{min}, \Delta_{max}] \subset [\tilde{\Delta}, \Delta_2(\epsilon)]$ for which we can prove the condensation. Clearly, this interval increases with decreasing ϵ. Alternatively, if for a given $\epsilon > 0$, one fixes $\Delta > \Delta_{min}$ (whereas in Theorem 1.1 we fixed μ), the bounds are only positive in a μ-interval, between a $\mu_{\min}$ and a μ_{max}. Hence we prove condensation only for a finite range of particle densities. It might come over as counter-intuitive that condensation would disappear again for high densities. However, we would like to notice that solvable models of Bose gases with this feature [21] are known in the literature. Remark also that nevertheless the upper bound μ_{max} increases with decreasing ϵ. These upper bounds on Δ or μ for a fixed ϵ are due to the fact that the terms of the highest order in μ and Δ appearing in the bounds (3.1) or (3.17) and (3.12) have negative sign and they are of the form:

$$-\frac{(\mu+\Delta)^2}{\Delta}\epsilon.$$

Hence, for smaller values of ϵ the domain, for which we can prove condensation, gets larger.

In Figure 2, the dependence of the lower bound on the temperature is indicated. Again,

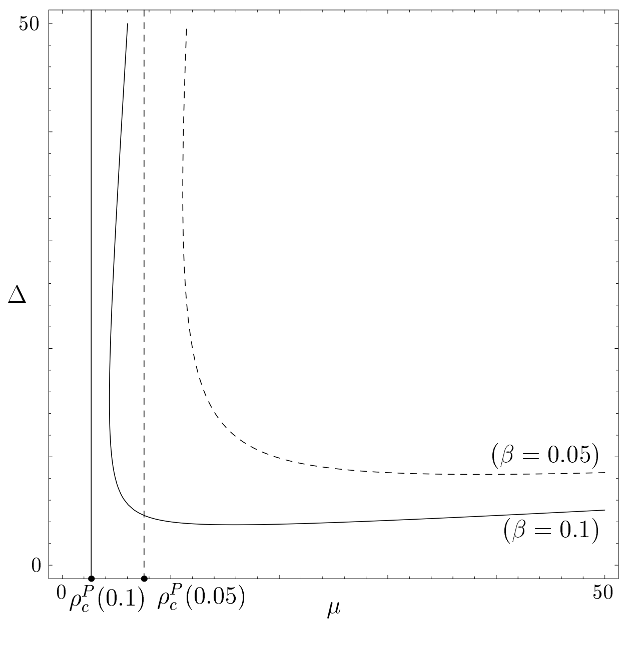

Figure 2: (μ, Δ)-graph with temperature dependence.

the lines shown on the (μ, Δ)-graph indicate domains where the lower bound (3.12) is positive, i.e. above each of these curves we have BEC. The dashed curve is the threshold for condensation calculated for inverse temperature $\beta = 0.05$, and the solid line is the threshold at inverse temperature $\beta = 0.1$, $\epsilon = 0.1$ in both cases. Clearly, for higher values of β, the condensation occurs for smaller gaps and for smaller values of μ, i.e. at lower densities.

Figure 3 is the (μ, β) phase-diagram computed for a choice of parameters $\Delta = g = v(0) = \hat{v}(0) = 1$. The dotted line is the critical line $\mu = g\hat{v}(0)\rho_c^P(\beta)$, i.e. only above

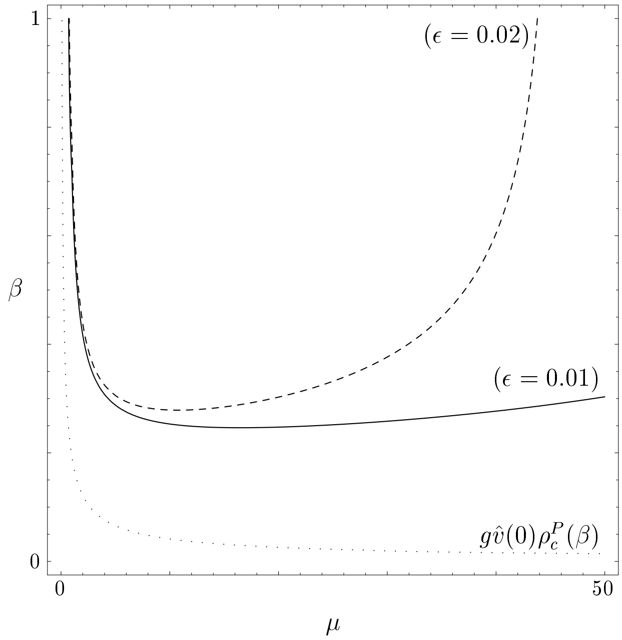

Figure 3: (μ, β) phase-diagram with ϵ-dependence.

this curve we have valid bounds on the condensate. The dashed curve is the threshold for condensation calculated for $\epsilon = 0.02$, and the solid line is the threshold at $\epsilon = 0.01$, calculated from the minorant (3.12).

On Figure 4 we present a family of thresholds for $\epsilon = 0.01$ as a function of the gap value Δ. As above, this family is calculated by equalising the lower bound (3.12) to zero.

Considering the high density (large μ) regime and taking ϵ of order μ^{-2}, the lower bound (3.12) can be written as

$$\frac{\mu}{\hat{v}(0)}\left(\frac{1}{g} - \frac{1}{\Delta}\left(\frac{v(0)}{2} + \hat{v}(0)\rho^P(\beta, -\Delta)\right)\right) + o(\mu),$$

which means that in order to have a positive lower bound for $\rho_{0,g}^{\Delta}(\beta, \mu)$, we need

$$\Delta > g\left(\frac{v(0)}{2} + \hat{v}(0)\rho^P(\beta, -\Delta)\right), \tag{4.1}$$

for high values of μ. This means that for non-zero interaction (1.3) ($g > 0$), there is a non-zero lower bound on the gap width. Notice that the minorant (4.1) is proportional to the coupling constant $g \geq 0$.

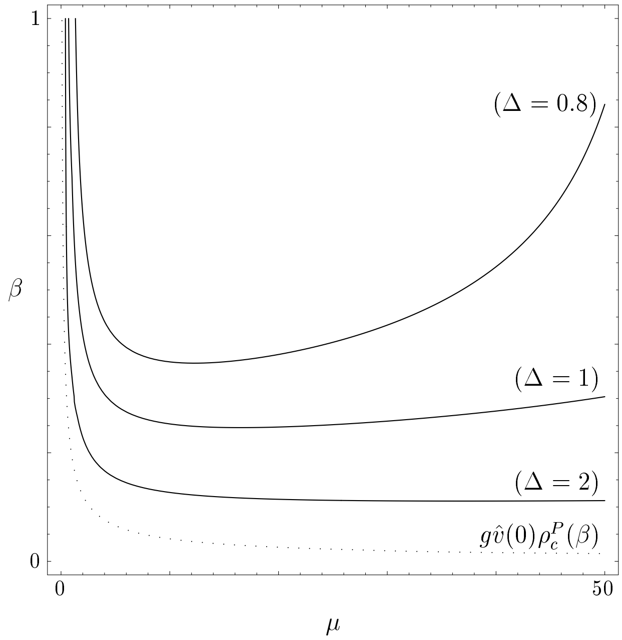

Figure 4: (μ, β) phase-diagram with Δ-dependence.

If we now choose for the two-body interaction a family of Van der Waals scaled pair potentials, i.e. we substitute

$$v(x-y) \mapsto \lambda^\nu v(\lambda(x-y)), \tag{4.2}$$

$\lambda > 0$, in the expression for the interaction term (1.3), then we find that for a fixed $\Delta > 0$, the condition (4.1) is satisfied if λ is chosen small enough and for low enough temperatures. This is easily seen as follows: substitution (4.2) amounts to substituting $\lambda^\nu v(0)$ for $v(0)$ and leaving $\hat{v}(0)$ in (4.1) unchanged. Therefore the r.h.s. of (4.1) can be made smaller than any $\Delta > 0$, by choosing λ small enough for large β such that $\rho^P(\beta, -\Delta)$ gets small. Hence, we recover the result of Buffet, de Smedt and Pulé [1] about the stability of Bose-Einstein condensation in the weakly interacting Bose gases with the Van der Waals scaled potentials and a non-zero one-particle spectral gap.

If $v(0)$ tends to infinity, then the condition (4.1) can not be satisfied for any finite gap. In this case our estimate becomes a triviality. This behaviour is compatible with the observation that for the hard core continuous Bose gas there is no condensation with a macroscopic occupation of any level [22].

Finally we remark that our results are for continuous homogeneous systems. The only assumptions we make are the gap in the one-particle excitations spectrum (1.2) and the superstability conditions on the pair-potential v (conditions (a) and (b)). Various other interesting exact results on Bose condensation are known, e.g. for Bose systems with a family of Van der Waals potentials [1,23], for models with truncated interactions [6,21,24], or for Bose lattice models with hard core interaction and at half-filling [25]. Only recently, a proof of BEC is found for the trapped interacting gases [2], i.e. for inhomogeneous systems.

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